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Feasibility study towards comparison of the $g^{(2)}(0)$ measurement in the visible range

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Abstract

This work reports on the pilot study, performed by INRIM, NPL and PTB, on the measurement of the $g^{(2)}(0)$ parameter in the visible spectral range of a test single-photon source based on a colour centre in diamond. The development of single-photon sources is of high interest to the metrology community as well as the burgeoning quantum technologies industry. Measurement of the $g^{(2)}(0)$ parameter plays a vital role in characterising and understanding single-photon emission. This comparison has been conducted by each partner individually using their own equipment at INRIM laboratories, which were responsible for the operation of the source.

Keywords: colour centers, single-photon source, metrology for quantum technologies

(Some figures may appear in colour only in the online journal)

1. Introduction

Single-photon sources (SPSs) [1–3], i.e. sources that are able to produce single photons on demand, can prove to be key elements for the development of quantum optical technologies. They will also be essential for providing metrological support for the development and commercialisation of these technologies, as well as for radiometry and photometry at the single-photon level. SPSs based on different physical systems (parametric down-conversion [4–11], quantum dots [12, 13], trapped ions [14], molecules [15] and colour centres in diamond [16–25]) and single-photon sensitive detectors [26, 27] and cameras [28] are widely available today as well as more complex equipment such as quantum key distribution systems [29, 30]. Despite several recent dedicated studies [31, 32], a standardized methodology for the characterization of SPSs has not emerged.

The typical parameter employed to test the properties of a SPS is the second order correlation function (or Glauber function) defined as

$$g^{(2)}(\tau = 0) = \left| \frac{\langle I(t)I(t + \tau) \rangle}{\langle I(t) \rangle^2} \right|_{\tau = 0},$$

where $I(t)$ is the intensity of the optical field. In the regime of low photon flux, this parameter has been shown to be substantially equivalent to the parameter $\alpha$ introduced by Grangier et al [33], which is experimentally measured as the ratio between the coincidence probability at the output of a Hanbury Brown and Twiss (HBT) interferometer [34], typically implemented by a 50:50 beam-splitter connected to two non-photon-number-resolving detectors, and the product of the click probabilities at the two detectors, i.e.:
\[ g^{(2)}(\tau = 0) \approx \alpha = \frac{P_C}{P_A P_B}, \]  

(2)

where \( P_C, P_A, P_B \) are, respectively, the coincidence and click probabilities at the outputs \( A, B \) of an HBT interferometer. This identity holds strictly for very low value of the click-intervals are used to estimate, respectively, background coincidences (a), true coincidences (b) and accidental coincidences (c).

Typical chronogram of the coincidence events registered with an HBT interferometer. The three identical highlighted time intervals are used to estimate, respectively, background coincidences (a), true coincidences (b) and accidental coincidences (c).

Figure 1 shows the typical chronogram of the behaviour of a pulsed SPS based on a nitrogen-vacancy centre excited in the pulsed regime, emitting single photons in the spectral range from 650 nm to 750 nm was used as a source.

An analogous effort to establish a proper procedure for the measurement of the \( g^{(2)} \) function of a telecom heralded SPS (in continuous regime) can be found in [35].

2. Measurement technique

With regards to equation (2), probabilities \( P_C, P_A, P_B \) are estimated as the ratio between the total number of the corresponding events versus the number of excitation pulses during the experiment, i.e. \( P_x = N_x/(\tau_{acq})(x = C, A, B) \), where \( R \) is the excitation rate and \( \tau_{acq} \) is the total acquisition time. The value of the measurand is independent from the total efficiencies \( (\eta_A, \eta_B) \) of individual channels (including detection and coupling efficiency), optical losses and splitting ratio since

\[ \alpha = \frac{\eta_A \eta_B P_C}{\eta_A P_A \eta_B P_B} = \frac{P_C}{P_A P_B}, \]  

(3)

The value of the parameter from the experimental data, corrected for the contribution of the background coincidences (due, for example, to stray light or residual excitation light), can be estimated as follows:

\[ \alpha = \frac{P_C - P_{Chg}}{(P_A - P_{Abg})(P_B - P_{Bbg})}, \]  

(4)

where \( P_{Chg}, P_{Abg}, P_{Bbg} \) are, respectively, the coincidence and click probabilities of background photons, calculated analogously to their counterparts \( P_C, P_A, P_B \).

Figure 1 shows the typical chronogram of the behaviour of a pulsed SPS obtained by sampling the coincidence events at the two outputs of an HBT interferometer. The coincidence probability has been estimated as the ratio between the total number of events in the chronogram falling in a fixed temporal...
window \( w \) around the central peak (showing antibunching, i.e. the ‘b’ interval in figure 1) and the total number of excitation pulses occurring in the acquisition time. The product \( P_{\text{APB}} \), corresponding to the probability of accidental coincidences, has been evaluated by integrating the events occurring in an equal interval around the subsequent peak (‘c’ interval in figure 1) not showing antibunching (always divided by the number of pulses). In fact, those coincidence events (amounting to \( N_\xi \)) are related to independent events (coincidences between single photons emitted after two subsequent laser pulses and detected by detector A and detector B respectively) and thus \( P_\xi - P_{\text{bg}} = (N_\xi - N_{\text{bg}})/(R \ast t_{\text{tacq}}) \).

The parameter to be estimated is thus:

\[
\alpha_{\text{exp}} = \frac{N_C - N_{\text{bg}}}{N_\xi - N_{\text{bg}}},
\]

where \( (N_i \) being the coincidence events sampled in the \( i \)th channel)

\[
N_C = \frac{k_w}{2} \sum_{i=-k_w/2}^{k_w/2} N_i,
\]

\[
N_\xi = \frac{T+k_w}{2} \sum_{i=T-k_w/2}^{T+k_w/2} N_i,
\]

\( k_w \) is the number of bins corresponding to the chosen coincidence window \( w \), \( N_{\text{bg}} \) is the estimated background due to spurious coincidences (the number of events in the ‘a’ interval in figure 1) and \( T \) is the excitation period (expressed in bins).

In figure 1 two backflash peaks [36–38] can be observed on either side of the central peak. These are due to secondary photon emission that arises from the avalanche of charge carriers that occurs in one of the two detectors in the HBT interferometer as a photon is absorbed and that are afterwards detected from the other detectors. To avoid overestimating \( \alpha \), these peaks must not be included in the coincidence window.

The presence of the backflash peaks prevented us to estimate \( P_A, P_B \) directly from the counts of the two detectors, since we were forced to consider a coincidence window smaller than the NV-center emission time window (of the order of tens of nanoseconds, i.e. at least three NV lifetimes). For this reason we estimated \( P_A, P_B \) consistently with the coincidences measured at time 0. The probability of observing a coincidence in the autocorrelation window around the peak at 400 ns can be underestimated by the presence of the coincidence counts between 0 and 400 ns, because of detectors and electronics dead-time. Due to the extremely low level of counts in this interval, we have estimated that this correction is negligible within the declared probability uncertainty.

3. Measurement facility

Figure 2 shows the experimental setup: a laser-scanning confocal microscope whose signal is split by a 50:50 beamsplitter and directed to two independent HBTs measurement systems performing the comparison. One of the measurement devices is held by the host institution (INRIM) while the other one is used, in turn, by the other two partners (NPL, PTB).
The INRIM facility was composed of a fused 50:50 fibre beam-splitter connected to two Excelitas SPCM-AQR-14-FC single-photon avalanche detectors (SPADs). Single and coincidence counts were sampled via ID Quantique ID800 time-to-digital converter (60 ps time resolution).

The NPL facility was composed of a fused 50:50 fibre beam-splitter connected to two Perkin-Elmer SPCM-AQR-14-FC single-photon avalanche detectors (SPADs). Coincidence counts were sampled via PicoQuant HydraHarp 400 multichannel picosecond event timer (1 ps time resolution).

The PTB facility was composed of a fused 50:50 fibre beam-splitter connected to two Excelitas SPCM-AQR-14-FC single-photon avalanche detectors (SPADs). Single and coincidence counts were sampled via PicoQuant HydraHarp 300 multichannel picosecond event timer (4 ps time resolution).

The detailed description of the sample fabrication and preparation is reported elsewhere [39].

### 4. Results

Each measurement consisted of 10 runs each of 500 s acquisition time. The total coupling rate, accounting for limited SPS quantum efficiency, collection angle, optical losses

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**Table 1.** Uncertainty Budget \((k = 2)\) associated with INRIM (in joint measurement with NPL).

<table>
<thead>
<tr>
<th>Quantity</th>
<th>Value</th>
<th>Standard unc.</th>
<th>Sens. coeff.</th>
<th>Unc. contribution</th>
</tr>
</thead>
<tbody>
<tr>
<td>(N_C)</td>
<td>1000</td>
<td>70</td>
<td>(1.5 \times 10^{-4})</td>
<td>(1 \times 10^{-2})</td>
</tr>
<tr>
<td>(N_\xi)</td>
<td>7400</td>
<td>900</td>
<td>(-9 \times 10^{-6})</td>
<td>(-1 \times 10^{-2})</td>
</tr>
<tr>
<td>(N_{BG})</td>
<td>560</td>
<td>30</td>
<td>(-1.4 \times 10^{-4})</td>
<td>(-3 \times 10^{-3})</td>
</tr>
<tr>
<td>(\alpha_{exp})</td>
<td>0.065</td>
<td>0.005</td>
<td>0.005</td>
<td>0.005</td>
</tr>
</tbody>
</table>

**Table 2.** Uncertainty Budget \((k = 2)\) associated with NPL measurement (in joint measurement with INRIM).

<table>
<thead>
<tr>
<th>Quantity</th>
<th>Value</th>
<th>Standard unc.</th>
<th>Sens. coeff.</th>
<th>Unc. contribution</th>
</tr>
</thead>
<tbody>
<tr>
<td>(N_C)</td>
<td>900</td>
<td>200</td>
<td>(2 \times 10^{-4})</td>
<td>(2 \times 10^{-2})</td>
</tr>
<tr>
<td>(N_\xi)</td>
<td>6000</td>
<td>2000</td>
<td>(-1 \times 10^{-5})</td>
<td>(-2 \times 10^{-2})</td>
</tr>
<tr>
<td>(N_{BG})</td>
<td>540</td>
<td>50</td>
<td>(-2 \times 10^{-4})</td>
<td>(-7 \times 10^{-3})</td>
</tr>
<tr>
<td>(\alpha_{exp})</td>
<td>0.068</td>
<td>0.005</td>
<td>0.005</td>
<td>0.005</td>
</tr>
</tbody>
</table>

**Table 3.** Uncertainty Budget \((k = 2)\) associated with INRIM measurement (in joint measurement with PTB).

<table>
<thead>
<tr>
<th>Quantity</th>
<th>Value</th>
<th>Standard unc.</th>
<th>Sens. coeff.</th>
<th>Unc. contribution</th>
</tr>
</thead>
<tbody>
<tr>
<td>(N_C)</td>
<td>800</td>
<td>100</td>
<td>(2 \times 10^{-4})</td>
<td>(3 \times 10^{-2})</td>
</tr>
<tr>
<td>(N_\xi)</td>
<td>5000</td>
<td>1000</td>
<td>(-2 \times 10^{-5})</td>
<td>(-2 \times 10^{-2})</td>
</tr>
<tr>
<td>(N_{BG})</td>
<td>380</td>
<td>30</td>
<td>(-2 \times 10^{-4})</td>
<td>(-6 \times 10^{-3})</td>
</tr>
<tr>
<td>(\alpha_{exp})</td>
<td>0.079</td>
<td>0.009</td>
<td>0.009</td>
<td>0.009</td>
</tr>
</tbody>
</table>
and detection efficiency (excluding the splitting ratio of the detector-tree), has been estimated as the ratio between the counting rate of the detectors (summing over all four of them) and the excitation rate, yielding \( \eta_{\text{TOT}} = (1.76 \pm 0.01)\% \). The coincidence window \( w \) considered for evaluating the reported \( \alpha \) was \( w = 16 \text{ ns} \). By repeating the analysis for different temporal widths \( w \) it was observed that the results were consistent as long as the backflash peaks were not included in the coincidence window (see figure 5). Figures 3 and 4 show the distributions of the \( \alpha_{\exp} \) values measured by each partner; the continuous line indicates the mean value and the dashed lines draw a 1-\( \sigma \) confidence band around the mean value. Tables 1–4 report the uncertainty budgets associated with the measurements. The summary of the results of the joint measurement is presented in table 5. We observe that individual measurement sessions (INRIM/NPL and INRIM/PTB) yield results that are extremely consistent. Mechanical instability in the coupling of the source may be the reason why the two sessions are not perfectly in agreement and the results of the INRIM/PTB joint measurements yield a slightly higher \( \alpha \) value (as well as greater associated uncertainty) with respect to the INRIM/NPL ones. In fact, the agreement in the INRIM/NPL measurements is better than indicated by figure 4 and the calculations, since the NPL measurements took longer than the INRIM measurements, the last two NPL measurements being performed after INRIM had completed its measurements. However, all values are compatible within the uncertainty

\[ u_c(\alpha_{\exp}) = \sqrt{\sum_x \left( \frac{\partial \alpha_{\exp}}{\partial N_x} \right)^2 u(N_x)^2 + 2 \sum_{x,y} \rho_{xy} \left( \frac{\partial \alpha_{\exp}}{\partial N_x} \right) \left( \frac{\partial \alpha_{\exp}}{\partial N_y} \right) u(N_x)u(N_y)} \]

\[(k = 2). \]

5. Dependence on the coincidence window

To prove that the estimation of \( \alpha \) is independent of the choice of the time interval of integration, we performed an analysis of the values of the measurand obtained by varying the coincidence window \( w \). The results are shown in figure 5, demonstrating that, as long as the backflash peaks are not included in the integration, the estimate is consistent independently of \( w \).

6. Conclusions

A pilot study on the characterization of a pulsed-pumped test SPS based on a NV centre in nanodiamonds was performed by INRIM, NPL and PTB and hosted by INRIM. This study will greatly benefit the single-photon metrology community, as

\[ \rho_{xy} = \frac{\langle N_xN_y \rangle - \langle N_x \rangle \langle N_y \rangle}{u(N_x)u(N_y)} \]
well as rapidly-growing quantum-technology-related industries. The main result of this study was the development of a standardized measurement technique as well as an uncertainty estimation procedure. The validity of the technique (system-independent and unaffected by the non-ideality of the apparatus) is demonstrated by the results obtained, yielding for all the participants estimated values of $g^{(2)}(0)$ that are compatible within the uncertainty ($k = 2$).

Acknowledgments

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Appendix. Lifetime estimation

The mean lifetime associated with the source has been estimated by numerically fitting the coincidence histograms (as in figure 1) via the single-exponential function [41–43]

$$f(\tau) = a + b \sum_{n=-\infty}^{\infty} \left(1 - \frac{\delta_{0n}}{c}\right) e^{-\frac{\tau - n\Delta\tau}{\tau}}, \quad (A.1)$$

where $a$ corresponds to the number of background coincidences, $b$ is a normalization factor, $\delta_{0n}$ is the Dirac Delta, $c$ is the number of excited emitters, $n$ is the excitation pulse number, $\Delta\tau$ is the excitation period and, finally, $d$ accounts for the lifetime (convoluted with the detectors’ jitter) of the center. Figure A1 shows the results of the lifetime estimation independently performed by the partners. Each value in the plot represents the mean of the results of 10 fits (one for each experimental run performed by one partner). Averaging the results, it is obtained the value $t_{\text{life}} = (15.34 \pm 0.08) \text{ ns}$.

References

[16] Aharonovich I, Englund D and Toth M 2016 Nat. Photon. 10 631
[26] Lolli L et al 2011 Int. J. Quantum Inf. 9 405
[34] Hanbury Brown R and Twiss R Q 1956 Nature 177 27
[38] Goetzberger A, McDonald B, Haitz R H and Scarlett R M 2004 J. Appl. Phys. 34 1591